

# Nuclear reactions, reactors

szerző: PGY

# Energy from fusion

When you join small atoms together, you can also get energy.

The Sun fuses hydrogen to make helium.

We're currently trying to fuse two isotopes of hydrogen - it's easier.



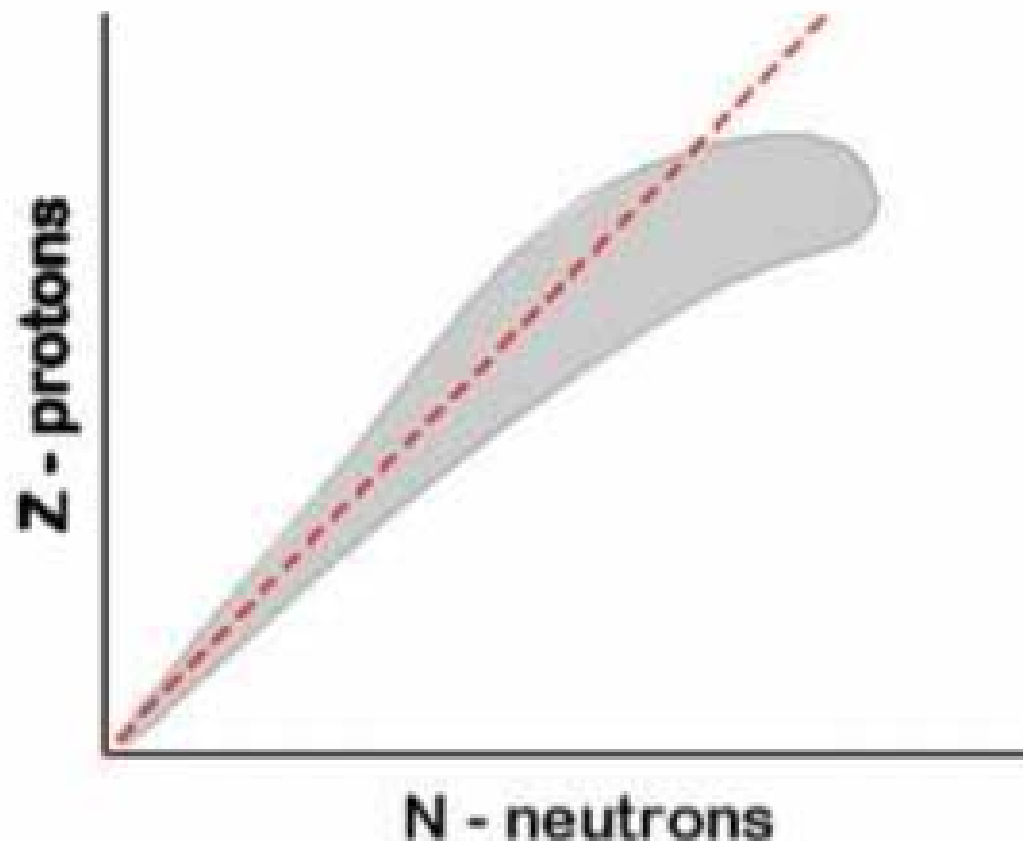
It will be great when it works. Fuel would come from the oceans - almost free.

# Chart of the Nuclides

**The radiochemist's version of the Periodic Table, listing all known nuclear isotopes.**

- **There are more than 2300 known nuclides and over 400 isomers.**
- **Only 287 isotopes are stable or naturally occurring radioactive forms.**
- **The chart lists this information along with other valuable data.**

# Chart of the Nuclides



The chart is arranged as a plot of neutron number Vs. atomic number.

Stable isotopes run at ~ 45° slope then slope down at  $Z = 20$  (Ca).

Due to the size, the chart is often split up into several pages.

## **Information obtained from the chart**

**Each chart 'box' lists appropriate types of physical data for a specific nuclide.**

**The style and color of the box also gives you a fair amount of information.**

**Since there is so much information, lets look at some examples.**

# Stable isotope

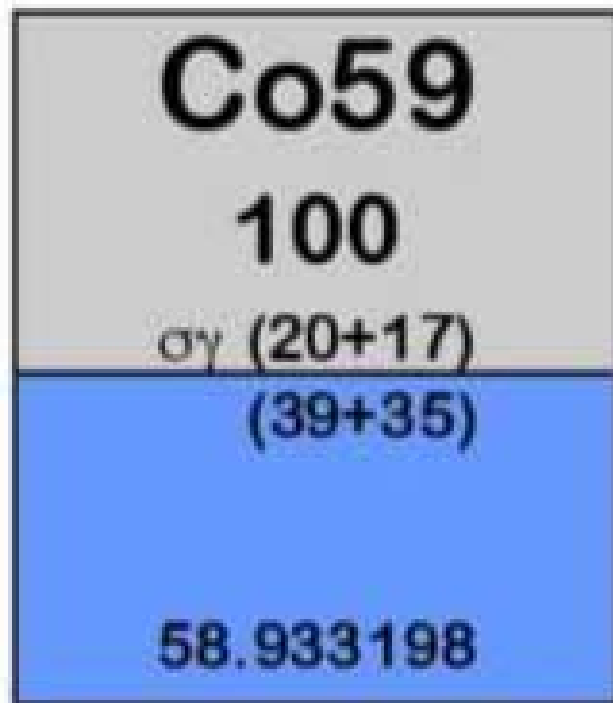
Stable isotopes are listed in gray boxes.

The diagram shows a gray rectangular box containing the following text from top to bottom: **C13**, **1.10**,  $\sigma_T$  1.4mb, 1.6mb, and **13.00335482**. Four arrows point from text labels on the right to the corresponding fields in the box: the top arrow points to **C13**, the second arrow points to **1.10**, the third arrow points to  $\sigma_T$  1.4mb, 1.6mb, and the bottom arrow points to **13.00335482**.

<b>C13</b>	Symbol & mass number
<b>1.10</b>	Percent abundance
$\sigma_T$ 1.4mb, 1.6mb	Thermal neutron and resonance cross sections.
<b>13.00335482</b>	Isotopic mass - C12 scale

Other information might be listed for other isotopes.

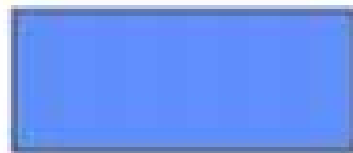
# Stable isotope



**This stable isotope also has a color code**

**A color on the bottom half of the box is used to indicate its relative large neutron cross section.**

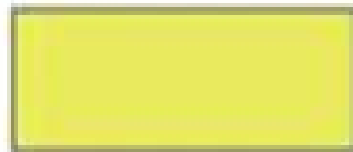
# Cross section color coding.



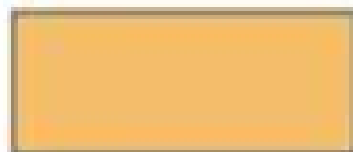
10 - 100 barns



100 - 500 barns



500 - 1000 barns



>1000 barns

**This is a measure of neutron absorption by a nuclide.**

**We'll cover in detail later when we discuss neutron activation.**



# Naturally occurring radioisotope

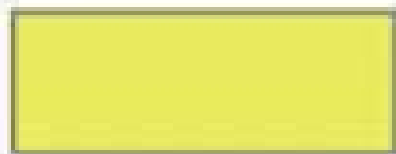
<b>H3</b>	
<b>12.3a</b>	← half life
$\beta^-$ .01860	← Mode and energy of decay (in MeV)
no $\gamma$	
$\sigma_\gamma < 6\mu\text{b}$	← Cross section
E .01860	← Beta disintegration energy in MeV.

Note: Color coding is also used to indicate relative half life values.

## Half life color coding.



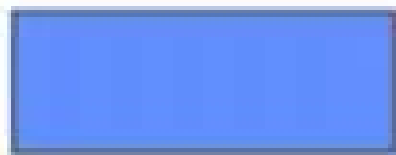
1 - 10 days



10 - 100 days



100 - 10 years



10 years - 10,000 years

Color coding of  
the top half of  
the box is used  
for half life

< 1 day and  
> 10,000 years  
are not color  
coded.

# Artificially radioactive

**Ca47**

**4.54 d**

$\beta^-$  .69, ...

$\gamma$  1.297, ...

**E 1.988**

**This example show  
the most common  
modes of decay.**

**Others may be  
present.**

# Symbols used in chart.

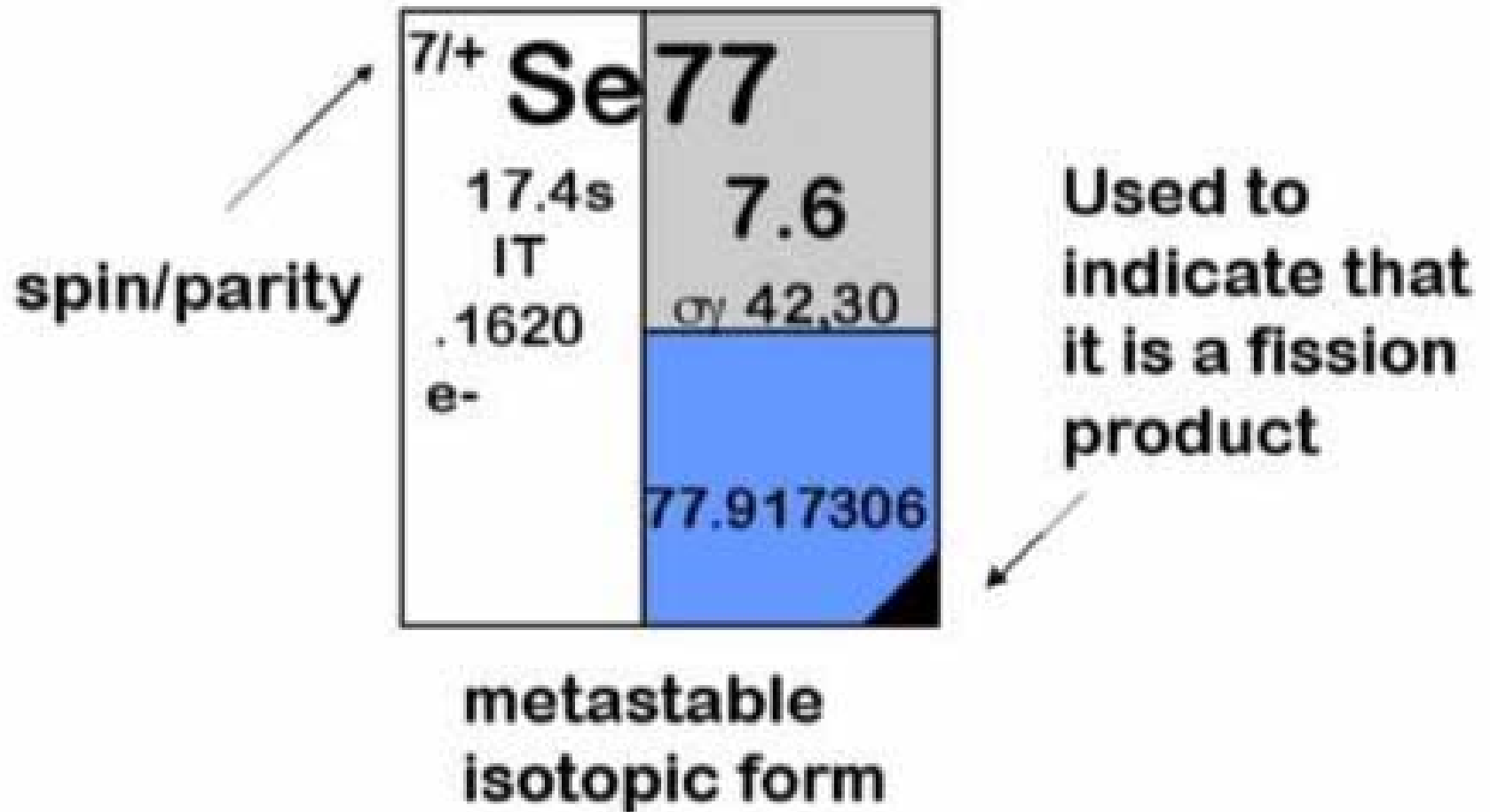
## Radiations and decay

$\alpha$	alpha
$\beta^-$	negative e-
$\beta^+$	positron
$\gamma$	gamma
n	neutron
p	proton
e	electron capture
IT	isomeric transition
SF	spontaneous fission
$\beta^-\beta^+$	double beta decay

## Time

$\mu\text{s}$	microseconds
ms	milliseconds
s	seconds
m	minutes
h	hours
d	days
a	years

# Two isomeric states - one stable



# Spin information

Each neutron and proton has an intrinsic angular momentum of  $1/2 h/2\pi$ .  $h$  is Plank's constant.

These combine with their orbital angular momentum to produce a resultant angular momentum called the **nuclear spin**.

Orbital angular momentum is always zero, so nuclear spin has an integer or half-odd-integer value depending on the nucleons.  
even - integer, odd -  $1/2$  odd integer

# Parity information

Mathematical formalism of quantum theory.

even parity (+)

odd parity (-)

**Example:** Al-27 has a value of 5/+

This is actually 5/2+ but the 2 has been omitted to improve readability.

The ground states of all even-even nuclides is 0+ so they are omitted from the chart.

## **So What!**

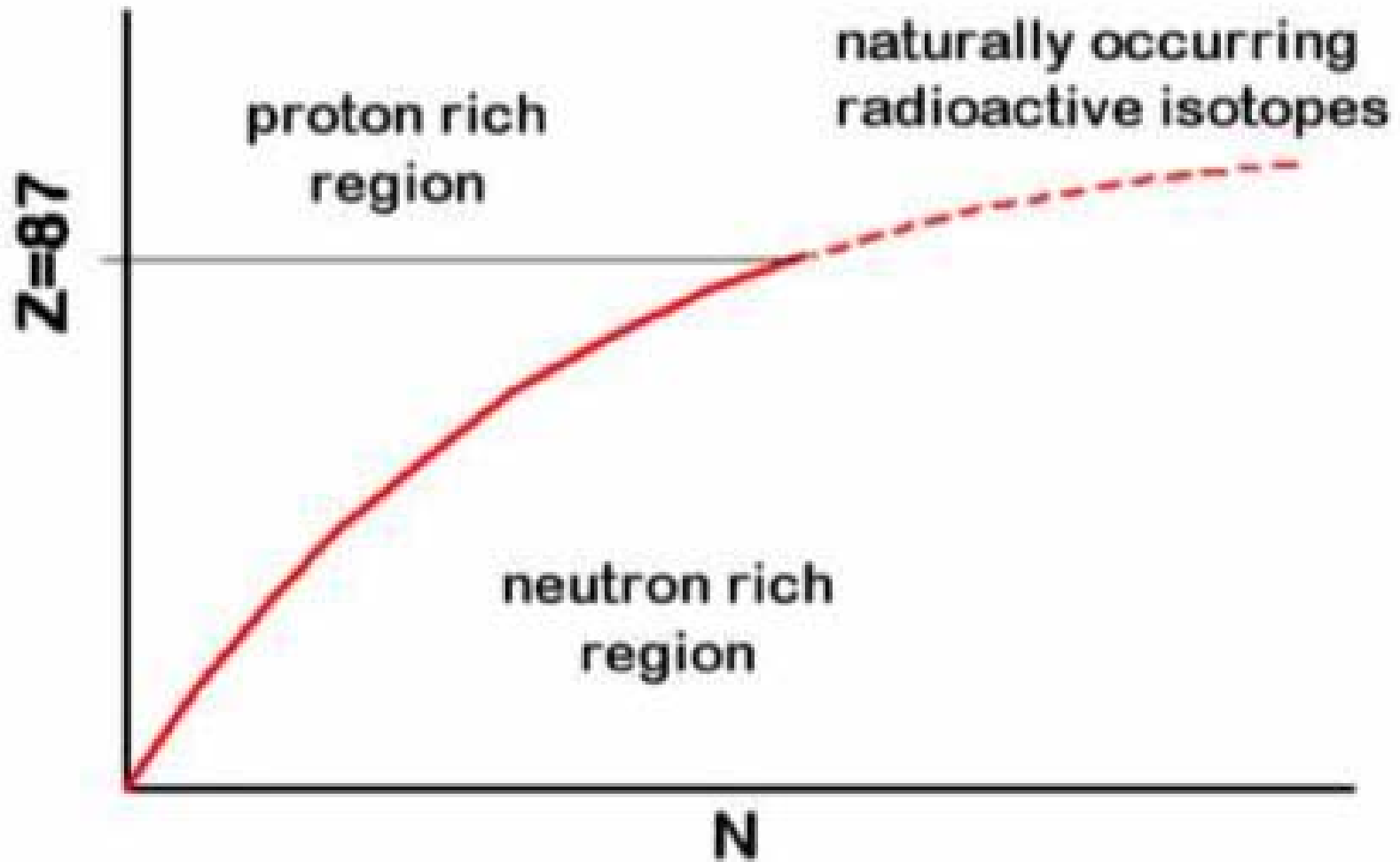
**A large angular momentum or spin change is required for gamma-ray transition between energy states of a nuclide**

**If the spin change is large and the energy of the transition is small, you end up with relatively long half lives for the transition.**

**As a result, you are able to see metastable states.**



# Decay trends and the chart



# Decay trends

## Naturally occurring radioactive isotopes

All are alpha emitters.

Too many nucleons.



This is a quick way to reduce the number by 4. Multiple steps may be required to complete the process - **decay series**.

# Decay trends

## Proton rich isotopes

Prefer  $\beta^+$  or  $\epsilon$  emission



In essence, we are converting a p to an n  
- trying to get back to the zone of stability.

# Decay trends

## Neutron rich isotopes

Prefer  $\beta^-$  emission



In essence, we are converting a neutron to a proton.

## Nuclear transmutations and the chart

Due to its arrangement, the chart can assist in determining what will form as a result of a nuclear reaction or decay.

### Symbols used:

n - neutron

$\alpha$  - alpha

p - proton

$\beta^-$  - beta

d - deuteron

$\beta^+$  - positron

t - tritium

$\epsilon$  - electron capture

# Nuclear transmutations

			${}^3\text{He}$ in	$\alpha$ in
	$\beta^-$ out	p in	d in	t in
	n out	original	n in	
t out	d out	p out	$\beta^+$ out $\epsilon$	
$\alpha$ out	${}^3\text{He}$ out			

**You can use the chart to rapidly determine what will be produced.**

# Nuclear transmutations

			${}^3\text{He}$ in	$\alpha$ in
	$\beta^-$ out	p in	d in	t in
	n out	original	n in	
t out	d out	p out	$\beta^-$ out $e^-$	
$\alpha$ out	${}^3\text{He}$ out			

O15	O16	O17	O17
N13	N14	N15	N15
C12	C13	C14	C14
B11	B12	C13	C13

What would result from the  $\beta^-$  decay of C12?

Inserting a deuterium into B12?

# Nuclear transmutations

Many nuclear reactions involve bombarding a nucleus with a particle to obtain a new species.

We use a form of 'nuclear short hand' to describe the reaction.

$X$  ( particles in , particles out )  $Y$

We omit gamma rays from the reaction since they don't affect the outcome.



# Nuclear transmutations

The (in,out) pattern can be used with the chart to predict what **could** be formed. This does not tell you if it will actually occur.

## Common bombardment patterns

(n, $\gamma$ )      (p, $\alpha$ )      ( $\gamma$ ,n)

(n,p)      ( $\gamma$ ,n)      (p, $\gamma$ )

# Nuclear transformations

	$\alpha, 3n$	$\alpha, 2n$	$\alpha, n$	
	$p, n$	$p, \gamma$ $d, n$	$\alpha, np$ $t, n$	
	$\gamma, n$ $n, 2n$	original $n, n$	$d, p$ $n, \gamma$	$t, p$
$p, \alpha$	$n, t$	$\gamma, p$	$n, p$	
	$n, \alpha$	$n, pd$		

This is a partial list of nuclear reactions.

## Example

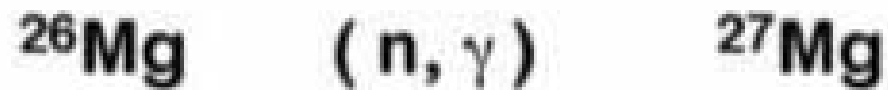
Starting with Mg-26, predict what will be formed by the following reactions.



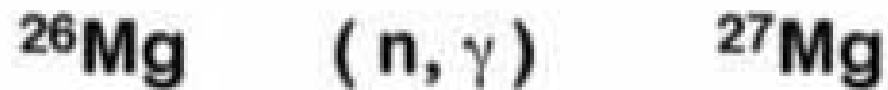
# Example



# Example



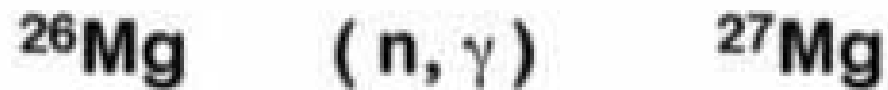
# Example



# Example

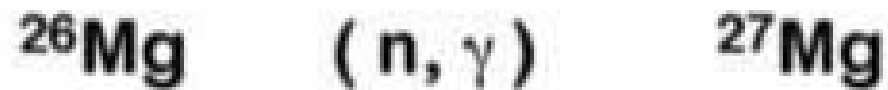


# Example





# Example



# Binding energy and the chart

**When protons and neutrons combine, mass is lost to energy - Binding energy.**

**If the nucleus was not more stable than the separate nucleons it would not have formed (or stay formed)**

**The binding energy of a nucleus can be found based on:**

$$E = mC^2$$

# Binding energy and the chart

A more useful version of the equation is

$$\Delta E = \Delta m c^2$$

where:

$\Delta E =$  the binding energy

$\Delta m =$  mass difference between the nucleus and the separate nucleons.

Since  $1 \text{ amu} = 931 \text{ MeV}$

$$\text{Binding energy} = \Delta m_{(\text{amu})} \times 931 \text{ MeV/amu}$$

## Example

Determine the binding energy of  $^{16}\text{O}$ .

First, look up the masses of  $^{16}\text{O}$ , p and n.

$^{16}\text{O}$	15.9949146 amu
n	1.00866497 amu
p	1.00782504 amu

## Example

Next, determine the total mass for the separate protons and neutrons.

$^{16}\text{O}$  - 8 protons and 8 neutrons

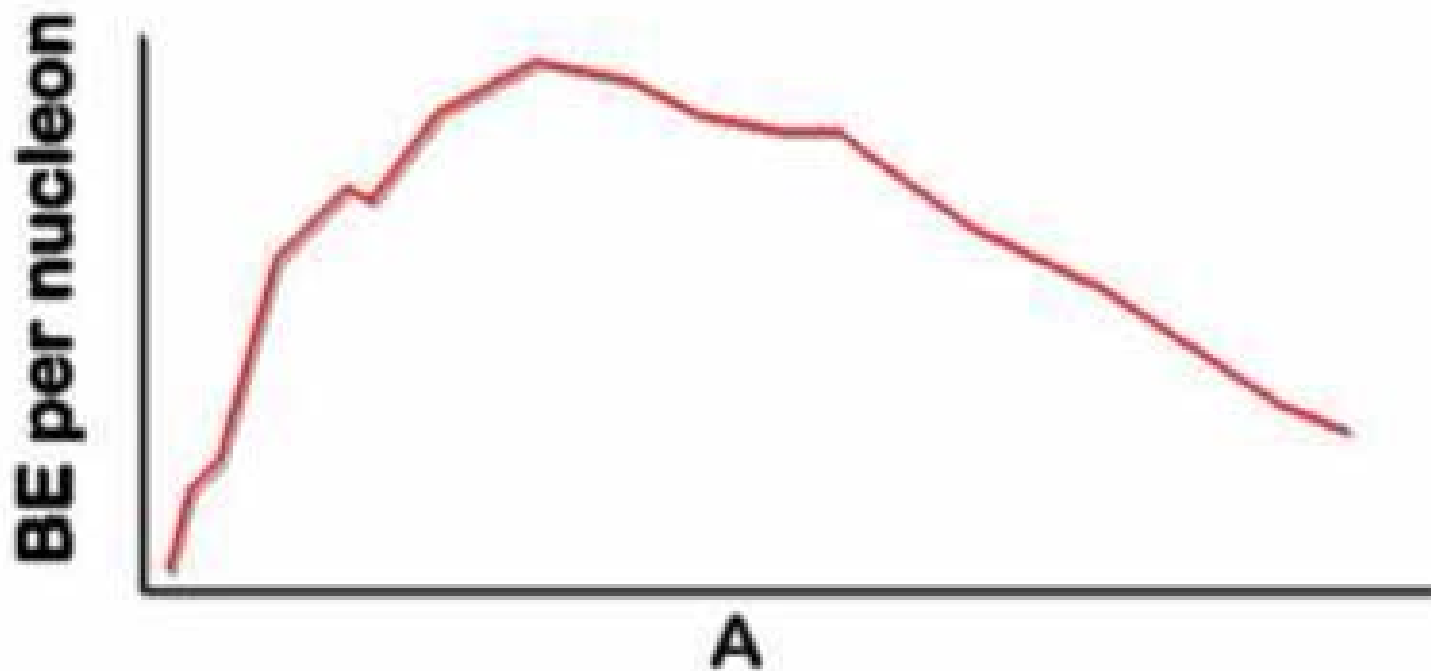
$$8 \text{ n} \quad 8 \times 1.00866497 = 8.0693197$$

$$8 \text{ p} \quad 8 \times 1.00782504 = 8.0620032$$

$$\text{Total} \quad 16.13192008$$

# Binding energy

We can calculate the binding energy for all of the stable isotopes and end up with the following plot.

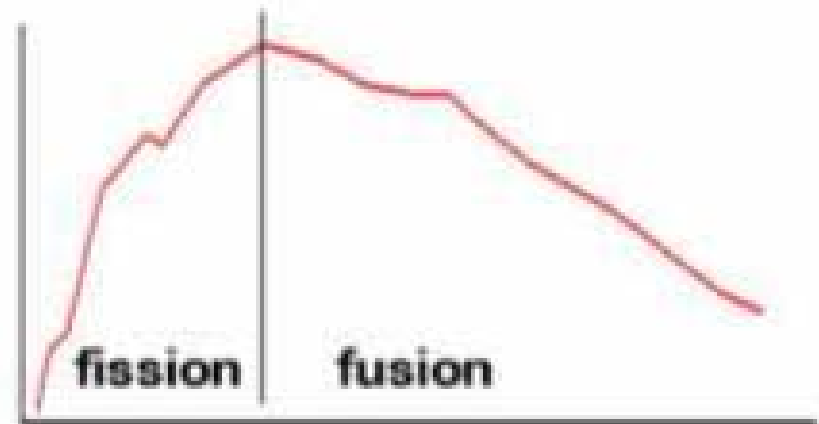


# Binding energy

As the total number of nucleons increases, we reach a point where we reach a maximum.

Higher mass nucleons are less stable.

This is why we can obtain energy from both fission and fusion and why alpha emission is common for heavier isotopes.



# Nuclear power

Power can be obtained two ways.

## **Fission**      Splitting atoms

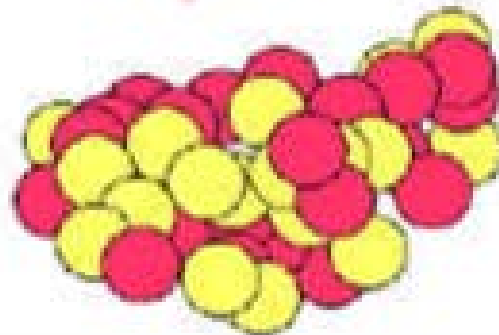
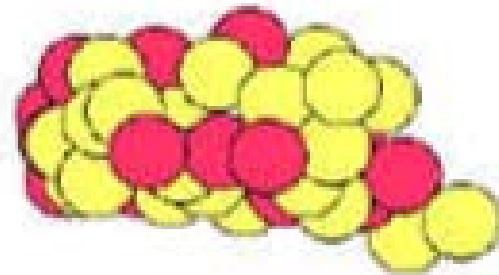
- Get energy if the nucleus is big.
- The smaller ones are more stable.
- What we do in nuclear reactors.

## **Fusion**      Joining atoms

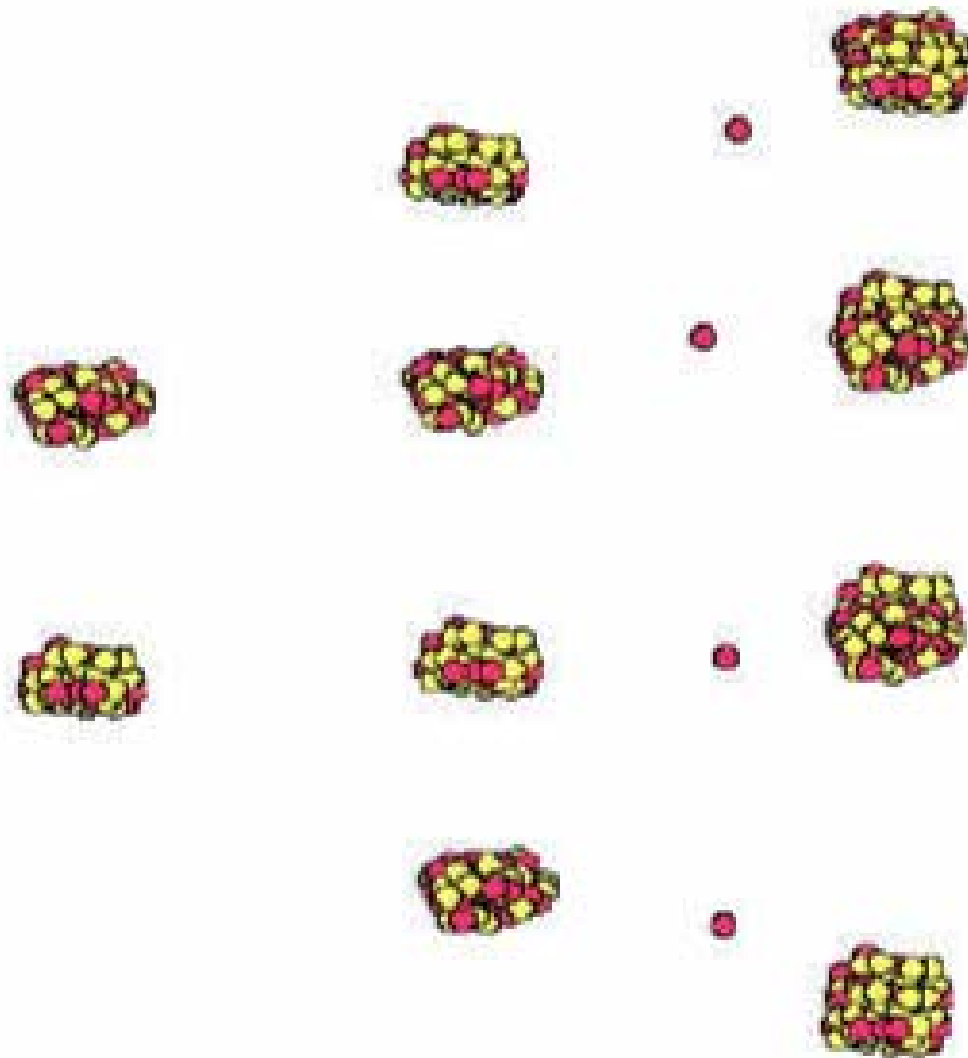
- Get energy if the nuclei are small.
- The larger one is more stable.
- This is how the sun works.



# Nuclear Fission



# Chain reaction



# Chain reactions

## Critical Reaction

When just enough fissions occur to keep the chain reaction reaction going.

- nuclear power

## Supercritical Reaction

When excess neutrons are produced and the rate of fission keeps increasing.

- nuclear bombs

# Energy from fission

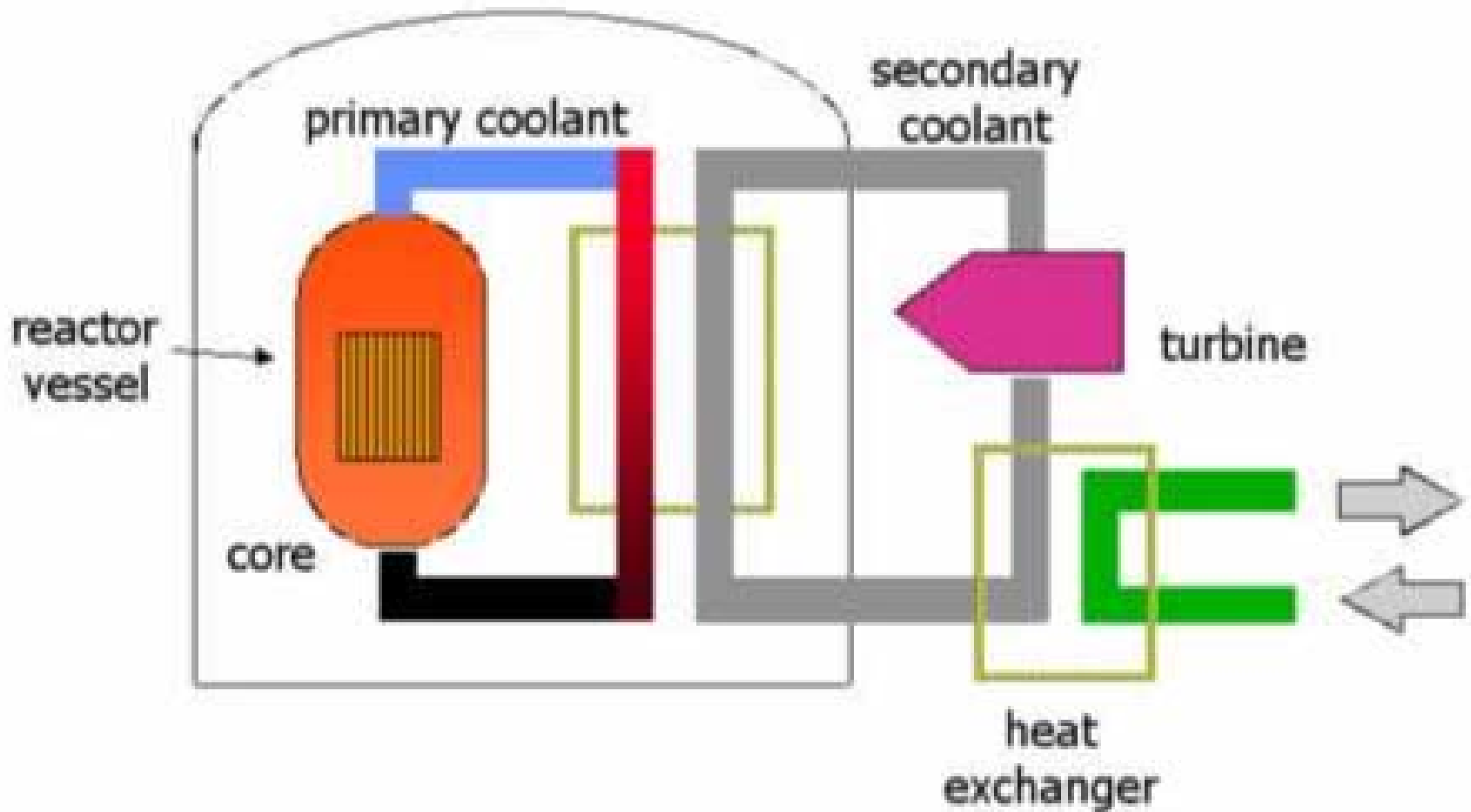
Uranium-235 is used as a 'fuel' in a reactor.  
One common reaction is



The energy produced by splitting one atom is approximately 200 million electron volts.

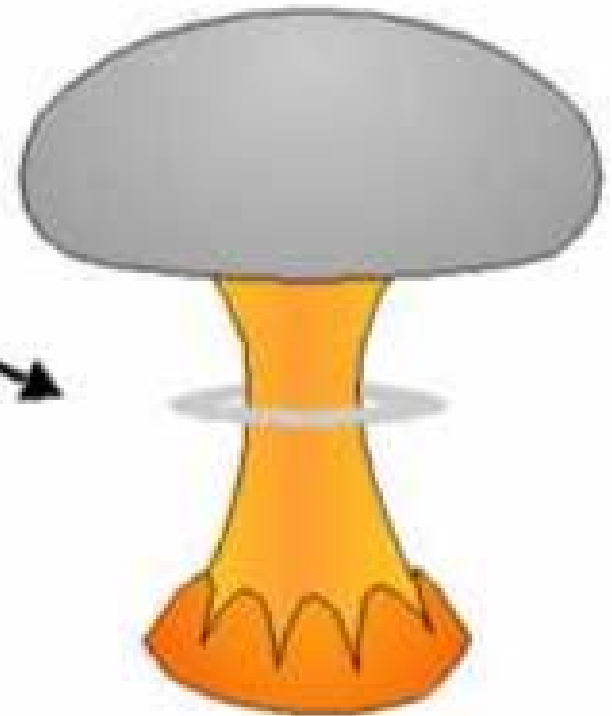
100 grams of  ${}^{235}\text{U}$  could produce as much energy as 80 trillion tons of TNT.

# Nuclear reactors



# Nuclear bombs

A conventional explosive is used to drive two sections of U-235 together.



This creates a supercritical mass.